

**PHYS 4410**  
**Quantum Mechanics 2**  
**Spring 2023**

**Lecture 31**

**Emission of electromagnetic radiation**

April 12

1 Today: quantum system interacting w/ EM waves

Dominant interactions:

$$V = -\vec{E} \cdot \vec{p}$$

↑ dipole moment.

$$e^{i\mathbf{k}\cdot\mathbf{r}} = e^{i\omega r/c} \rightarrow 1$$

non-relativistic:  $c \rightarrow \infty$

$$\vec{p} = -e\vec{r} \quad (\text{for hydrogen atom})$$

Light = EM wave: 
$$\vec{E} = E_0 \hat{\epsilon} \left[ e^{i(\vec{k}\cdot\vec{r} - \omega t)} + e^{-i(\vec{k}\cdot\vec{r} - \omega t)} \right]$$

↑ polarization (unit vector)

$$\omega = c|\vec{k}|$$

Via Fermi's Golden Rule:

start in  $|i\rangle$ , transition rate to  $|f\rangle$ :

$$R_{i \rightarrow f} = \frac{2\pi}{\hbar^2} \delta(\omega_{fi} - \omega) |\langle f | E_0 \hat{\epsilon} \cdot \vec{p} e^{i\vec{k}\cdot\vec{r}} | i \rangle|^2$$

"electric dipole approx"

$$\omega_{fi} = \frac{E_f - E_i}{\hbar}$$

$$= \frac{2\pi}{\hbar^2} E_0^2 \delta(\omega_{fi} - \omega) |\langle f | \hat{\epsilon} \cdot \vec{p} | i \rangle|^2$$

2 Incoherent light:

$$E_0 \hat{\epsilon} e^{i\vec{k}\cdot\vec{r} - i\omega t} \rightarrow \sum_{2\hat{\epsilon}} \int d^3k \underbrace{E_0(k)}_{\text{"random phase"}} \hat{\epsilon}_k e^{i\vec{k}\cdot\vec{r} - i\omega t}$$

$$c_f^{(1)} \sim -\frac{i}{\hbar} \int_0^t ds e^{i\omega_f s} \sum_{2\hat{\epsilon}} \int d^3k e^{-i\vec{k}\cdot\vec{r}} E_0(k) \hat{\epsilon} \cdot \langle f | \hat{\rho} | i \rangle$$

$$P_{i \rightarrow f} = |c_f^{(1)}|^2 \sim \int d^3k d^3k_* \underbrace{E_0(k) E_0(k_*)^*}_{\text{average over?}} \dots$$

$|E_0(k)| \cdot |E_0(k_*)^*| \cdot e^{i\vec{k}\cdot\vec{r} - i\vec{k}_*\cdot\vec{r}}$   
 except if  $\vec{k} = \vec{k}_*$

$$P_{i \rightarrow f} \sim \int d^3k |E_0(k)|^2 \dots$$

$$R_{i \rightarrow f} = \int d^3k R_{i \rightarrow f}(k)$$

[rates add,  
not amplitudes  
for incoherent]

3  $R_{i \rightarrow f} = \frac{2\pi}{\hbar^2} \int d^3k \delta(\omega_{fi} - c|\vec{k}|) \underbrace{|\langle f | \hat{\epsilon} \cdot \vec{p} | i \rangle|^2}_{\text{average over } k, \hat{\epsilon}: \text{isotropic average}} \underbrace{|E_0(k)|^2}_{\frac{\rho(k)}{2\epsilon_0} \leftarrow \text{energy dens.}}$

$$= \frac{1}{3} |\langle f | p_x | i \rangle|^2 + \frac{1}{3} |\langle f | p_y | i \rangle|^2 + \frac{1}{3} |\langle f | p_z | i \rangle|^2$$

$$= \frac{1}{3} |\langle f | \vec{p} | i \rangle|^2 \quad (\text{implicit dot product})$$

$$\rho(k) = \frac{2}{(2\pi)^3} \frac{\hbar \omega}{e^{\hbar \omega / k_B T} - 1} \quad (\omega = c|\vec{k}|)$$

$$\int d^3k \delta(\omega_{fi} - c|\vec{k}|) \frac{\rho(k)}{2\epsilon_0} \longrightarrow \int d\omega \frac{\hbar \omega^3}{\pi^2 c^3 (e^{\hbar \omega / k_B T} - 1)} \dots$$

$$= \int d\omega \rho(\omega)$$

rate of photons  
absorbed ( $E_f > E_i$ )

$E_f < E_i$ : stim.  
emission.

$$\rightarrow R_{i \rightarrow f} = \frac{\pi}{3\epsilon_0 \hbar^2} \rho(\omega_{fi}) \underbrace{|\langle f | \vec{p} | i \rangle|^2}$$

4  $H_0 = \frac{p^2}{2m} + \frac{1}{2}m\Omega^2 x^2$  and  $P = -ex$

If our initial state  $|i\rangle = |3\rangle$ , allowed  $|f\rangle$ ?

Write  $x = \sqrt{\frac{\hbar}{2m\Omega}}(a + a^\dagger)$ .

$\langle f | P | i \rangle = -e \sqrt{\frac{\hbar}{2m\Omega}} \langle f | a + a^\dagger | 3 \rangle$

What is frequency of absorb/emit photon?

$\langle f | \sqrt{3} | 2 \rangle + \langle f | \sqrt{4} | 4 \rangle$

$f=2$  or  $f=4$ .

$\downarrow$   
 $\frac{E_4 - E_3}{\hbar} = \Omega$

$-\Omega = \frac{(\frac{5}{2} - \frac{3}{2})\hbar\Omega}{\hbar} = \frac{E_2 - E_3}{\hbar}$

Rate:  $R_{3 \rightarrow 2} = \frac{\pi}{3\epsilon_0 \hbar^2} \rho(\Omega) \frac{|\langle 2 | a | 3 \rangle|^2}{3} \left( \frac{e^2 \hbar}{2m\Omega} \right)$

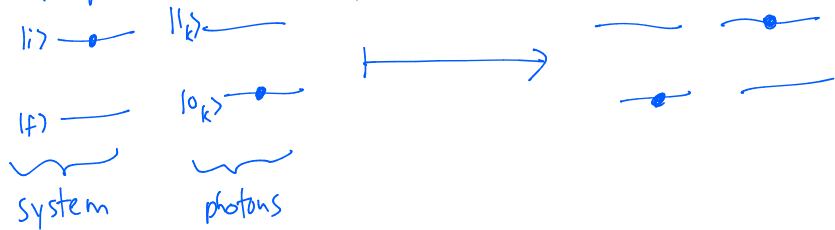
Suppose  $|i\rangle = |2\rangle$  and  $|f\rangle = |3\rangle \dots$  runaway heating...

$R_{2 \rightarrow 3} = R_{3 \rightarrow 2}$



5 Einstein: spontaneous emission  $|3\rangle \rightarrow |2\rangle$ ,  
 McIntyre: stat mech.

Only possible if EM quantize:



Energy:

$$E_i + \frac{1}{2} \hbar \omega_k$$

$$E_f + \frac{3}{2} \hbar \omega_k$$

$V =$  interaction b/w'n photon + system

degenerate:  $E_f - E_i + \hbar \omega_k = 0.$

6 Can generalize time-independent PT ...

$$R_{i \rightarrow f}^{se} = \frac{2\pi}{\hbar^2} \int d^3k \left| \langle f | \langle 1_k | (-\vec{E} \cdot \vec{p}) | i \rangle \otimes | 0_k \rangle \right|^2 \delta(\omega_{fi} - ck)$$

(A<sub>fi</sub>)

$$\sim \frac{1}{3} |\langle f | \vec{p} | i \rangle|^2 \quad \text{QED: } \sqrt{\frac{\hbar \omega}{8\pi^3 \epsilon_0}}$$

$$R_{i \rightarrow f}^{se} = \frac{2\pi}{\hbar^2} \int \frac{4\pi \omega^2 d\omega}{c^3} \delta(\omega_{fi} - \omega) \cdot \frac{1}{3} |\langle f | \vec{p} | i \rangle|^2 \cdot \frac{\hbar \omega}{8\pi^3 \epsilon_0}$$

$$= \frac{\omega_{fi}^3}{3\epsilon_0 \hbar \pi c^3} |\langle f | \vec{p} | i \rangle|^2$$

rate of spontaneous emission