

**PHYS 5210**  
**Graduate Classical Mechanics**  
**Fall 2022**

**Lecture 15**  
**The Klein-Gordon equation**

September 26

**1** Use effective field theory to deduce the minimal theory for a single relativistic field.

Goal: one field  $\phi$ ,  $\mathcal{L}$  invariant under trans./Lorentz

Build  $\mathcal{L}$  out of "invariant building blocks":

Ex 1: B.B. =  $\phi$       $\phi(x) \rightarrow \phi(x+\epsilon)$  isn't invariant?

but.  $S = \int d^4x \mathcal{L}_1(\phi(x)) \mapsto \int d^4(x+\epsilon) \mathcal{L}(\phi(x+\epsilon))$   
 $\underbrace{d^4x}_{dt dx dy dz} \quad \underbrace{(\dots)}_{x'} \quad \underbrace{d^4(x+\epsilon)}_{x'} = \int d^4x' \mathcal{L}(\phi(x'))$

Ex. 2: B.B. =  $\partial_\mu \phi \partial^\mu \phi = \partial_\mu \phi \partial_\nu \phi \eta^{\mu\nu}$

$[\partial_\mu \phi \rightarrow \Lambda_\mu^\nu \partial_\nu \phi \neq \partial_\mu \phi \text{ is not B.B.}]$

Our most general  $\mathcal{L}$  [w/o 2-der. on  $\phi$ ]:  $\mathcal{L}_2(\phi, \partial_\mu \phi \partial^\mu \phi)$

Suppose  $\phi$  small ( $\phi \approx 0$ ). Taylor expand  $\mathcal{L}_2$ :

Klein-Gordon

$$\mathcal{L}_2 = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{m^2}{2} \phi^2$$

[ $\phi$ -linear term dropped for stability]

2

What are the equations of motion for a Klein-Gordon field? Interpret them in terms of relativistic particle dynamics.

K-G theory:  $\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{m^2}{2} \phi^2$

EOMs:  $0 = \frac{\delta S}{\delta \phi} = \underbrace{-m^2 \phi}_{\frac{\partial \mathcal{L}}{\partial \phi}} + \underbrace{\partial_\mu \partial^\mu \phi}_{\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)}}$

$$\begin{aligned} \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} &= \frac{\partial}{\partial (\partial_\mu \phi)} \left( -\frac{1}{2} \eta^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi \right) \\ &= -\frac{1}{2} \eta^{\alpha\beta} (\delta_{\alpha}^{\mu} \partial_\beta \phi + \delta_{\beta}^{\mu} \partial_\alpha \phi) \\ &= -\partial^\mu \phi \end{aligned}$$

For relativistic physics,  
work in units w/  $c=1$ , ( $\hbar=1$ )

$$m^2 \phi = \partial_\mu \partial^\mu \phi = -\partial_t^2 \phi + \partial_x^2 \phi + \partial_y^2 \phi + \partial_z^2 \phi$$

Plane wave solutions!

$$\phi = e^{ikx - i\omega t}$$

$$m^2 e^{ikx - i\omega t} = -(-i\omega)^2 e^{ikx - i\omega t} + (ik)^2 e^{ikx - i\omega t}$$

$\omega^2 = k^2 + m^2$

Most general:

$$\phi = \int d^3k \left[ A(\mathbf{k}) e^{ikx + i\sqrt{k^2 + m^2}t} + B(\mathbf{k}) e^{ikx - i\sqrt{k^2 + m^2}t} \right]$$

$$(mc)^2 + (pc)^2 = E^2$$

QM:  $\omega = E/\hbar$ ,  $k = p/\hbar$

Naive:  $\phi$  as wave function?  
Not quite.

3 What is the stress-energy tensor?

K-G theory: invariant under  $x^\mu \rightarrow x^\mu + \epsilon^\mu = x^\mu + \epsilon^\nu \delta_\nu^\mu$

Goal: conserved stress-energy tensor  $T^{\mu\nu}$ . ( $\partial_\mu T^{\mu\nu} = 0$ )

Noether's Thm: if  $x^\mu \rightarrow x^\mu + \epsilon^\nu Y_\nu^\mu$  leaves  $S$  inv.

$$(\partial_\mu T^{\mu\nu} = 0) \quad J_\nu^\mu = \mathcal{L} Y_\nu^\mu - \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \partial_\nu \phi$$

Translations:  $Y_\nu^\mu = \delta_\nu^\mu$  plug in K-G  $\mathcal{L}$

$$T^\mu_\nu = \mathcal{L} \delta_\nu^\mu - \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \partial_\nu \phi = \partial^\mu \phi \partial_\nu \phi - \delta_\nu^\mu \frac{1}{2} (m^2 \phi^2 + \partial_\rho \phi \partial^\rho \phi)$$

Use  $\eta$  to raise  $\nu$ :

$$\eta^{\nu\rho} T^\mu_\nu = \boxed{T^{\mu\rho} = \partial^\mu \phi \partial^\rho \phi - \frac{1}{2} \eta^{\mu\rho} (m^2 \phi^2 + \partial_\sigma \phi \partial^\sigma \phi)}$$

4 Why is  $T^{\mu\nu} = T^{\nu\mu}$ ?

$$T^{\mu\nu} = \partial^\mu \phi \partial^\nu \phi - \frac{i}{2} \eta^{\mu\nu} (m^2 \phi^2 + \partial_\sigma \phi \partial^\sigma \phi) = T^{\nu\mu}$$

$$T^t i \underset{\substack{\uparrow \\ \text{spatial index}}}{=} T^i t, \quad \text{up to factors of } c: (\text{momentum dens.})_i = i\text{-comp. of energy cur.}$$

Why? Noether's Thm applied to Lorentz!

$$\Lambda^\mu_\rho \Lambda^\nu_\sigma \eta^{\rho\sigma} = \eta^{\mu\nu}$$

Plug in  $\Lambda = \delta + \epsilon$ :

$$\Lambda^\mu_\nu = \delta^\mu_\nu + \epsilon^\mu_\nu$$

$$x^\mu \rightarrow x^\mu + \epsilon^{\mu\nu} x_\nu$$

$$\epsilon^{\mu\nu} x_\nu = \frac{1}{2} \delta^\mu_\rho \underbrace{[\epsilon^{\rho\nu} x_\nu - \epsilon^{\nu\rho} x_\nu]}_{\text{trans?}}$$

↓ conserved current:

$$J^\mu(\epsilon) = \frac{1}{2} T^\mu_\rho [\epsilon^{\rho\nu} x_\nu - \epsilon^{\nu\rho} x_\nu] = \underline{J^\mu_{\rho\nu}} \epsilon^{\nu\rho}$$

where  $\partial_\mu J^\mu_{\nu\rho} = 0$

$$\eta^{\mu\nu} = \delta^\mu_\rho \delta^\nu_\sigma \eta^{\rho\sigma} + \epsilon^\mu_\rho \eta^{\rho\nu} + \epsilon^\nu_\sigma \eta^{\mu\sigma}$$

$$0 = \epsilon^{\mu\nu} + \epsilon^{\nu\mu}$$

$$\partial_\mu J^\mu_{\nu\rho} = \partial_\mu \frac{1}{2} (T^\mu_\rho x_\nu - T^\mu_\nu x_\rho)$$

$$\left( \cancel{\partial_\mu T^\mu_\rho} \right) x_\nu + T^\mu_\rho \partial_\mu x_\nu - \left( \cancel{\partial_\mu T^\mu_\nu} \right) x_\rho - T^\mu_\nu \partial_\mu x_\rho$$

$$T^\mu_\rho \eta_{\mu\nu} - T_{\nu\rho}$$

$$\rightarrow = T_{\nu\rho} - T_{\rho\nu} = 0.$$