

PHYS 5210
Graduate Classical Mechanics
Fall 2023

Lecture 13
Lagrangian field theory

September 27

To day: field theory describe systems of $N \gg 1$ particles w/ spatial locality.

Example: vibration of solid (phonons)

[assume identical mass m , spring constant k]

Lagrangian:
$$L = \sum_{j=-\infty}^{\infty} \left[\underbrace{\frac{1}{2} m \dot{u}_j^2}_{\text{term } j} - \frac{1}{2} k (u_{j+1} - u_j)^2 \right]$$

EOMs:
$$\frac{\delta S}{\delta u_j} = 0 = -k \left[\underbrace{(u_j - u_{j-1})}_{\text{term } j} + \underbrace{(u_j - u_{j+1})}_{\text{term } j-1} \right] - m \ddot{u}_j$$

or
$$m \ddot{u}_j = -k \underbrace{[2u_j - u_{j+1} - u_{j-1}]}$$

locality in space (& time)

Usually: only interested in long-distance physics...

Solid object: ~ 1 m vs. atomic distance $\sim 10^{-10}$ m

vibrating on scales $\gg 10^{-10}$ m: $u_j \approx u_{j+1}$

$u_j \approx u(j)$ [a continuous function?]
reasonable if " $\frac{du}{dj}$ " \ll u

Then: $2u_j - u_{j+1} - u_{j-1} \approx 2u - \left[u + \frac{\partial u}{\partial j} + \frac{1}{2} \frac{\partial^2 u}{\partial j^2} + \dots \right] - \left[u - \frac{\partial u}{\partial j} + \frac{1}{2} \frac{\partial^2 u}{\partial j^2} - \dots \right]$
 $\rightarrow - \frac{\partial^2 u}{\partial j^2}$

Convert to physical units: $a \sim 10^{-10}$ m (interatomic distance)

$j = \frac{x}{a}$, or $\frac{\partial}{\partial j} \rightarrow a \frac{\partial}{\partial x}$

Think of $u_j(t) \rightarrow u(x, t)$ as a field, function of spatial coords (x) as well as t .

EOMs: $m \frac{\partial^2 u}{\partial t^2} = ka^2 \frac{\partial^2 u}{\partial x^2}$ (one-dimensional mass density)

divide by a : $\frac{m}{a} = \rho$, $\kappa = ka$

$\rho \frac{\partial^2 u}{\partial t^2} = \kappa \frac{\partial^2 u}{\partial x^2} \rightarrow$ sound waves!

Can we derive PDE directly from Lagrangian
effective field theory (EFT)?

Yes! But need to generalize Lagrangian mechanics to fields.

Idea:
$$L = \sum_{j=-\infty}^{\infty} \left[\frac{a\rho}{2} u_j^2 - \frac{\kappa}{2a} (u_{j+1} - u_j)^2 \right]$$

$$\downarrow$$

$$\approx \sum_{j=-\infty}^{\infty} a \left[\frac{\rho}{2} \left(\frac{\partial u}{\partial t}(a, j, t) \right)^2 - \frac{\kappa}{2a^2} \left(a \frac{\partial u}{\partial x} + \frac{a^2}{2} \frac{\partial^2 u}{\partial x^2} + \dots \right)^2 \right]$$

Interpret as Riemann sum for integral...

$$\sum_{j=-\infty}^{\infty} a (1 + a \dots) \rightarrow \int_{-\infty}^{\infty} dx (1 + \underline{0})$$

$$L = \int_{-\infty}^{\infty} dx \mathcal{L} \quad \text{where } \mathcal{L} = \frac{\rho}{2} \left(\frac{\partial u}{\partial t} \right)^2 - \frac{\kappa}{2} \left(\frac{\partial u}{\partial x} \right)^2$$

(Lagrangian density)

and action
$$S = \int dt L \rightarrow \int dt dx \mathcal{L}$$

In EFT, directly write down \mathcal{L} based on symmetries, locality.

E.g. solids: symmetry of
$$u(x, t) \rightarrow u(x, t) + \varepsilon.$$

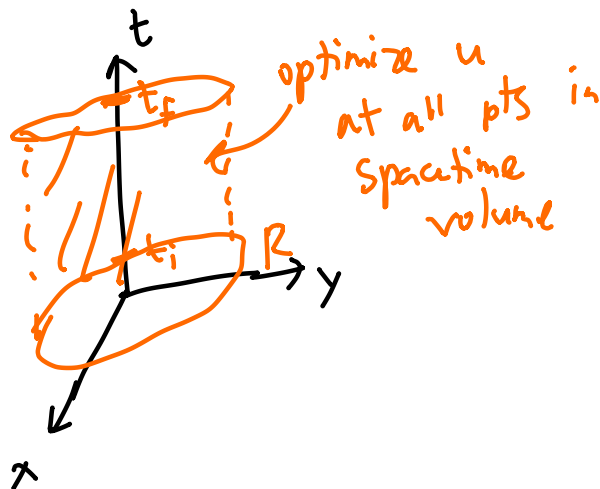
Before invariant BBs... generalize Euler-Lagrange to fields:

How to evaluate $\frac{\delta S}{\delta u(x)}$?

For ease of notation: $(x, t, \dots) \rightarrow x^\mu$
 (any add'l spatial dims)

NOT implying relativistic invariance.

$$S[u] = \int_{t_i}^{t_f} dt \int_R dV \mathcal{L}$$



As before, define functional derivative as:

$$\left. \frac{d}{d\varepsilon} S[\bar{u} + \varepsilon \hat{u}] \right|_{\varepsilon=0} = \int dt dV \hat{u} \left. \frac{\delta S}{\delta u} \right|_{\bar{u}}$$

POLA: $\frac{\delta S}{\delta u} = 0$ on physical "trajectories" of field.

$$S[u] = \int \underbrace{d^D x}_{\substack{\text{"}\prod d x^\mu\text{"} \\ \mu}} \mathcal{L}(u, \underline{\partial_\mu u}, \dots)$$

$$S[\bar{u} + \varepsilon \hat{u}] = \int d^D x \mathcal{L}(\bar{u} + \varepsilon \hat{u}, \partial_\mu \bar{u} + \varepsilon \partial_\mu \hat{u})$$

$$\frac{dS}{d\varepsilon} = \int d^D x \left[\hat{u} \frac{\partial \mathcal{L}}{\partial u} + \left(\sum_\mu \right) (\partial_\mu \hat{u}) \frac{\partial \mathcal{L}}{\partial (\partial_\mu u)} \right] \leftarrow \text{integrate by parts}$$

$$\hookrightarrow \int d^{D-1} x \hat{u} \frac{\partial \mathcal{L}}{\partial (\partial_\mu u)} \hat{u}^\mu - \int d^D x \hat{u} \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu u)}$$

$$= \int d^D x \partial_\mu \left(\hat{u} \frac{\partial \mathcal{L}}{\partial (\partial_\mu u)} \right)$$

Choose boundary conditions on \hat{u} so $\hat{u} = 0$ on spacetime bdy

$$\frac{dS}{d\varepsilon} = \int d^D x \hat{u} \left[\frac{\partial \mathcal{L}}{\partial u} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu u)} \right]$$

$$= \frac{\delta S}{\delta u} \left[= \frac{\partial \mathcal{L}}{\partial u} - \frac{\partial}{\partial t} \left(\frac{\partial \mathcal{L}}{\partial (\frac{\partial u}{\partial t})} \right) - \frac{\partial}{\partial x} \left(\frac{\partial \mathcal{L}}{\partial (\frac{\partial u}{\partial x})} \right) \right]$$

Natural generalization of $\frac{\delta S}{\delta u} = \frac{\partial \mathcal{L}}{\partial u} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial (\frac{du}{dt})}$.

$$\mathcal{L} = \frac{\rho}{2} (\partial_t u)^2 - \frac{\kappa}{2} (\partial_x u)^2$$

[shorthand: $\partial_t u = \frac{du}{dt}$, etc.]

↓

$$\frac{\delta S}{\delta u} = 0 = \cancel{\frac{\partial \mathcal{L}}{\partial u}} - \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial (\partial_t u)} - \frac{\partial}{\partial x} \frac{\partial \mathcal{L}}{\partial (\partial_x u)}$$

$$= -\rho \partial_t^2 u + \kappa \partial_x^2 u$$

or: $\rho \partial_t^2 u = \kappa \partial_x^2 u$.