

**PHYS 5210**  
**Graduate Classical Mechanics**  
**Fall 2023**

**Lecture 15**

**Noether's Theorem in field theory**

October 2

Lagrangian field theory:

$$S[\phi^a(x^\mu)] = \int d^D x \mathcal{L}(\phi^a, \partial_\mu \phi^a, \pi^\mu)$$

$\uparrow$   $\uparrow$  spacetime coords  
 $n$  fields ( $a=1, \dots, n$ )    ( $\mu=1, \dots, D$ )

$S$  made out of invariant building blocks under symmetries.

For particles: Noether Thm: continuous symmetry  $\rightarrow$  conserved quantity.

$$\frac{dQ}{dt} = 0.$$

on phys. traj.

Generalization to fields:

cont. sym  $\rightarrow$  conserved current:

$$0 = \partial_\mu J^\mu = \frac{\partial J^t}{\partial t} + \frac{\partial J^x}{\partial x} + \dots$$

If we restrict to  $D=1$ :  $0 = \partial_\mu J^\mu = \partial_t Q$

$$\begin{aligned}
 [\partial_\mu x^\nu = \delta_\mu^\nu] \\
 \downarrow \\
 \partial_\mu x^\rho = \delta_\mu^\nu \eta_{\nu\sigma} \\
 = \eta_{\mu\rho}
 \end{aligned}$$

Derivation of following is similar to lecture 3:

Define continuous symmetry: infinitesimal transformation...

- $\mathcal{L}(x^\mu, \phi^a) = \mathcal{L}(\tilde{x}^\mu, \tilde{\phi}^a) + \epsilon \partial_\mu K^\mu$  [ $\epsilon$  infinitesimal]
- $\tilde{x}^\mu = x^\mu + \epsilon X^\mu$
- $\tilde{\phi}^a = \phi^a(x^\mu) + \epsilon \varphi^a(x^\mu) = \phi^a(\tilde{x}^\mu) + \epsilon [\varphi^a(\tilde{x}^\mu) - X^\nu \partial_\nu \phi^a(\tilde{x}^\mu)] + \dots$

where  $S[\tilde{\phi}] = S[\phi]$ . (action is invariant)

$S$  invariant if:

$$0 = \mathcal{L} \partial_\mu X^\mu + \partial_\mu K^\mu + X^\mu \frac{\partial \mathcal{L}}{\partial x^\mu} + \varphi^a \frac{\partial \mathcal{L}}{\partial \phi^a} + (\partial_\nu \varphi^a - \partial_\nu X^\mu \partial_\mu \varphi^a) \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi^a)}$$

If satisfied  $\rightarrow$  continuous symmetry.

Noether's Thm: on solutions to EOM,  $\partial_\mu J^\mu = 0$  where

$$J^\mu = X^\mu \mathcal{L} + K^\mu + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^a)} (\varphi^a - X^\nu \partial_\nu \phi^a)$$

$\partial_\mu J^\mu = 0$  implies there's  $Q$  with  $\frac{dQ}{dt}$ :

$$Q = \int \underbrace{dx dy dz}_{\text{NOT time}} J^t \quad \text{density of "Q"}$$

$$\frac{dQ}{dt} = \int dx \dots \frac{\partial J^t}{\partial t} = \int dx \dots \left( -\frac{\partial J^x}{\partial x} - \dots \right) = 0$$

if spatial bdy doesn't add/remove charge (almost always)

Example 1: massless Klein-Gordon theory

$$S[\phi] = -\frac{1}{2} \int d^D x \partial_\mu \phi \partial^\mu \phi \quad (\underline{m=0})$$

Shift symmetry:  $\phi \rightarrow \phi + \epsilon$ .

$$\psi = 1, \quad \chi^\mu = 0, \quad K^\mu = 0$$

Conserved current:  $J^{\mu} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = -\partial^\mu \phi$

raise

lower

$$\partial_\mu J^\mu = -\partial_\mu \partial^\mu \phi = 0$$

$$\hookrightarrow \frac{\delta S}{\delta \phi} = \frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = 0.$$

Example 2: spacetime translation symmetry (almost always!)

$\mathbb{D}$  symmetries  $\leftarrow \chi^\mu = \text{const.}$

$$\left[ \frac{\partial \mathcal{L}}{\partial x^\mu} = 0 \right]$$

$$\textcircled{1} \begin{matrix} x^1 = 1 \\ x^2 = 0 \\ \vdots \end{matrix}$$

$$\textcircled{2} \begin{matrix} x^1 = 0 \\ x^2 = 0 \\ \vdots \end{matrix}$$

...

$$\textcircled{D} \begin{matrix} x^1 = 0 \\ \vdots \\ x^D = 1 \end{matrix}$$

choice of  $v = \text{symmetry}$

Package together:  $\chi^\mu = \overset{\text{const.}}{\downarrow} \epsilon^\mu \rightarrow \chi^\mu = \delta_\nu^\mu \epsilon^\nu$

$\uparrow$   $\mu$  index = label of spatial coords

Symmetry holds for any  $\epsilon^\nu \rightarrow \chi^\mu = \delta_\nu^\mu (\times \epsilon^\nu)$

$$K^\mu = 0$$

$$\psi^\alpha = 0.$$

Noether Thm:  $J^\mu_\nu(\xi) = \underbrace{J^\mu_\nu}_{\xi^\nu}$

$$J^\mu_\nu = \delta^\mu_\nu \mathcal{L} - \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \delta^\rho_\nu \partial_\rho \phi = \delta^\mu_\nu \mathcal{L} - \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \partial_\nu \phi$$

$J^\mu_\nu$  is conserved  $[\partial_\mu J^\mu_\nu = 0]$  for  $\alpha^\nu$  translation.

In relativistic theories:

pick one  
"word" 1

pick one  
"word" 2

$$J^\mu_\nu \eta^{\nu\rho} \mapsto T^{\mu\rho} = \begin{matrix} \text{energy-momentum} \\ \text{stress} \\ \text{stress-energy} \end{matrix}$$

current/tensor  
tensor

all combos OK.

$$\partial_\mu T^{\mu\rho} = 0, \quad T^{\mu\rho} = \eta^{\mu\rho} \mathcal{L} - \partial^\rho \phi \frac{\partial \mathcal{L}}{\partial(\partial_\rho \phi)}$$

Conservation of energy ( $X^t = 1$ ): ( $\mu = t$ )

$$E = \int \underbrace{d^3x}_{\text{space coords only}} T^{tt}$$

$T^{tt}$   
 $\uparrow$  density of conserved  
 $\nwarrow$  time-component  $\nu$  (energy cons)

Klein-Gordon:  $\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{m^2}{2} \phi^2 = \frac{1}{2} (\partial_t \phi)^2 - \frac{1}{2} (\partial_x \phi)^2 - \dots$

$$T^{tt} = -\mathcal{L} - \left( \eta^{tt} \partial_t \phi \right) \frac{\partial \mathcal{L}}{\partial(\partial_t \phi)}$$

$$= \underbrace{\frac{1}{2} (\partial_t \phi)^2}_{\text{"kinetic" energy}} + \underbrace{\frac{1}{2} (\partial_x \phi)^2 + \dots}_{\text{"strain" potential energy}} + \underbrace{\frac{1}{2} m^2 \phi^2}_{\text{"harmonic osc. potential"}}$$