

PHYS 5210
Graduate Classical Mechanics
Fall 2024

Lecture 6
Relativistic particles

September 9

Goal: today: action for relativistic particle of mass m
next: charged w/ EM fields

Step 1: identify symmetries:

① translation:

$$\begin{aligned} t &\rightarrow t + \epsilon_t \\ x &\rightarrow x + \epsilon_x \\ y &\rightarrow y + \epsilon_y \\ z &\rightarrow z + \epsilon_z \end{aligned}$$

4 independent translation generators

② rotation:

$$\begin{pmatrix} x \\ y \end{pmatrix} \rightarrow \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix}$$

rotate around x, y, z -axis

③ boost:

$$\begin{pmatrix} ct \\ x \end{pmatrix} \rightarrow \begin{pmatrix} \gamma & -\gamma\beta \\ -\gamma\beta & \gamma \end{pmatrix} \begin{pmatrix} ct \\ x \end{pmatrix}$$

move in the x, y, z -direction

where $\beta = \frac{v}{c}$ and $\gamma = (1 - \beta^2)^{-1/2}$.

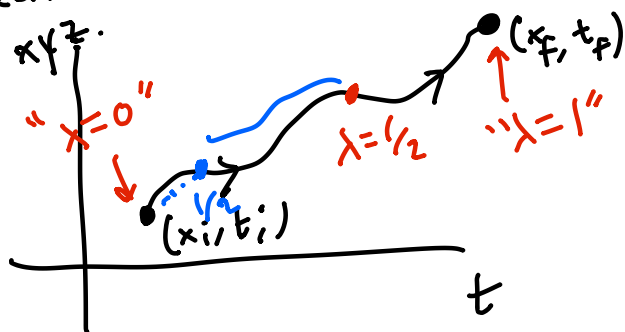
How many continuous symmetries? $4 + 3 + 3 = 10$. completely fix $S!$

Note: t & x on (almost) equal footing

↳ but $S = \int dt L(x, \dot{x}, \dots, t)$

$\nwarrow \frac{dx}{dt}$

Idea: introduce "fake time λ ":



Arbitrary choice of λ
so long as if $\lambda_1 < \lambda_2$,
 $t(\lambda_1) < t(\lambda_2)$

"time reparameterization"

λ is unphysical: $\lambda \rightarrow f(\lambda)$ so long as f is invertible

Goal: $S[t(\lambda), x(\lambda), y(\lambda), z(\lambda)] = \int d\lambda L(x^\mu, \frac{dx^\mu}{d\lambda}, \dots)$

Notation $x^\mu = \begin{pmatrix} x^t \\ x \\ y \\ z \end{pmatrix}$

Work in units where $c=1$.

Step 2: Identify invariant building blocks for L .
(Ignore reparameterization)

① Translations: $x^\mu \rightarrow x^\mu + \epsilon^\mu$ ($\epsilon_t, \epsilon_x, \epsilon_y, \epsilon_z$)

↳ this implies

$\frac{dx^\mu}{d\lambda}$ is invariant ... $\frac{\partial L}{\partial x^\mu} = 0$. (lec 2 & 3)

② & ③ Rotation & boost = Lorentz group

$\begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix}$

$\begin{pmatrix} t \\ x \end{pmatrix} = \begin{pmatrix} \gamma & -\gamma\beta \\ -\gamma\beta & \gamma \end{pmatrix} \begin{pmatrix} t \\ x \end{pmatrix}$

Under just z -rotation: t, z are invariant
 invariant $(x, y) = x^2 + y^2$



Add x/y -rotation: $t, x^2 + y^2 + z^2$ invariant

combined invariant under boost

$$-\Delta t^2 + (\Delta x^2 + \Delta y^2 + \Delta z^2) = -\Delta \tau^2$$

↑ "proper time"

Write as $\eta_{\mu\nu} \Delta x^\mu \Delta x^\nu$

repeat index: $\sum_{\mu\nu}$

(for relativistic: one upper/one lower)

$$\text{where } \eta_{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

In relativity, 'use $\eta_{\mu\nu}$ to lower indices':

$$x^\mu = \begin{pmatrix} t \\ x \\ y \\ z \end{pmatrix}, \quad \text{then} \quad \eta_{\mu\nu} x^\nu = x_\mu = \begin{pmatrix} -t \\ x \\ y \\ z \end{pmatrix}$$

So $\eta_{\mu\nu} x^\mu x^\nu = x_\mu x^\mu$. Analogue of vector's length.

Alternate picture: start w/ postulate that $x_\mu x^\mu$ is invariant under Lorentz

Useful: objects traveling @ $v < c$ have $\Delta x_\mu \Delta x^\mu = 0$ in all frames, all observers agree on speed of light

Look for infinitesimal coordinate changes: $x^\mu \rightarrow \Lambda^\mu_\nu x^\nu$

$$\Lambda^\mu_\nu = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} + \epsilon^\mu_\nu$$

↓ ϵ^μ_ν
 $\leftarrow \delta^\mu_\nu$

↑
const. matrix

$$\text{w/ } \eta_{\mu\nu} x^\mu x^\nu = \eta_{\mu\nu} (\Lambda^\mu_\alpha x^\alpha) (\Lambda^\nu_\beta x^\beta)$$

Taylor expand to first order in ϵ :

$$\eta_{\alpha\beta} = \eta_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta = \eta_{\mu\nu} (\delta^\mu_\alpha + \epsilon^\mu_\alpha) (\delta^\nu_\beta + \epsilon^\nu_\beta)$$

$$\rightarrow \text{first order: } 0 = \eta_{\mu\beta} \epsilon^\mu_\alpha + \eta_{\alpha\nu} \epsilon^\nu_\beta$$

$$-\epsilon^\mu_\alpha \eta_{\mu\beta} = \eta_{\alpha\nu} \epsilon^\nu_\beta \quad (-\epsilon^T \eta = \eta \epsilon)$$

$$\epsilon^\mu_\nu = \begin{pmatrix} 0 & \beta_x & \beta_y & \beta_z \\ \beta_x & 0 & \theta_z & -\theta_y \\ \beta_y & -\theta_z & 0 & \theta_x \\ \beta_z & \theta_y & -\theta_x & 0 \end{pmatrix}$$

3 boosts

3 rotations

ext-component:

$$+\epsilon_{tx} \eta_{xt} = \eta_{xx} \epsilon_{xt}$$

Conclude: under translation ① $\frac{dx^\mu}{d\lambda}$ invariant, as $\frac{d^2 x^\mu}{d\lambda^2}, \dots$

② & ③: invariants have all indices contracted!

$$\frac{dx^\mu}{d\lambda} \eta_{\mu\nu} \frac{dx^\nu}{d\lambda} = \frac{dx^\mu}{d\lambda} \frac{dx_\mu}{d\lambda} \quad \checkmark$$

Now invoke "fake reparameterization": $\lambda \rightarrow f(\lambda)$

$$S = \int d\lambda \quad L\left(\frac{dx^\mu}{d\lambda} \frac{dx_\mu}{d\lambda}\right)$$

$$\int df(\lambda) = f'(\lambda) d\lambda$$

$$\frac{dx^\mu}{d\lambda} \rightarrow \frac{dx^\mu}{df(\lambda)} = \frac{dx^\mu}{d\lambda} \frac{d\lambda}{df(\lambda)} = \frac{1}{f'} \frac{dx^\mu}{d\lambda}$$

S mustn't change for any $f(\lambda)$. So const.

$$S = \int d\lambda \quad \cancel{f'(\lambda)} \frac{1}{\cancel{f'(\lambda)}} \sqrt{-\frac{dx^\mu}{d\lambda} \frac{dx_\mu}{d\lambda}} \quad (-m)$$

Now, choose $\lambda = t$:

$$S = -m \int dt \sqrt{-\left[\left(\frac{dt}{dt}\right)^2 + \dot{x}^2 + \dot{y}^2 + \dot{z}^2\right]}$$

$$= -m \int dt \sqrt{1 - \frac{\dot{x}^2 + \dot{y}^2 + \dot{z}^2}{c^2}}$$

Restore c : ($c \rightarrow \infty$)

Taylor expand in c : $L = -mc^2 + \frac{m}{2}(\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$